

Numerical Simulation of Lake Breezes: Dependence on Location of Cloud Shadows on Land Surfaces

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Submitted: xx December 2023

Revised: xx xxxx xxxx

For publication in: *J. Adv. Model. Earth Syst.*

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Abstract

Simple parametrizations are developed that attempt to align high spatial-resolution distributions of surface solar irradiances predicted by 1D Independent Column Approximation (ICA) radiative transfer (RT) models with those from 3D RT models. Diffuse irradiances are smoothed to approximate horizontal diffusion below clouds, while direct-beam-induced cloud shadows are repositioned properly across the surface. Pearson correlation coefficients for global irradiances from a full 3D RT model and a 1D-ICA are ~ 0.2 and ~ 0.85 at low and high Sun, respectively. These values typically increase to ~ 0.87 and ~ 0.97 for the parametrization developed here.

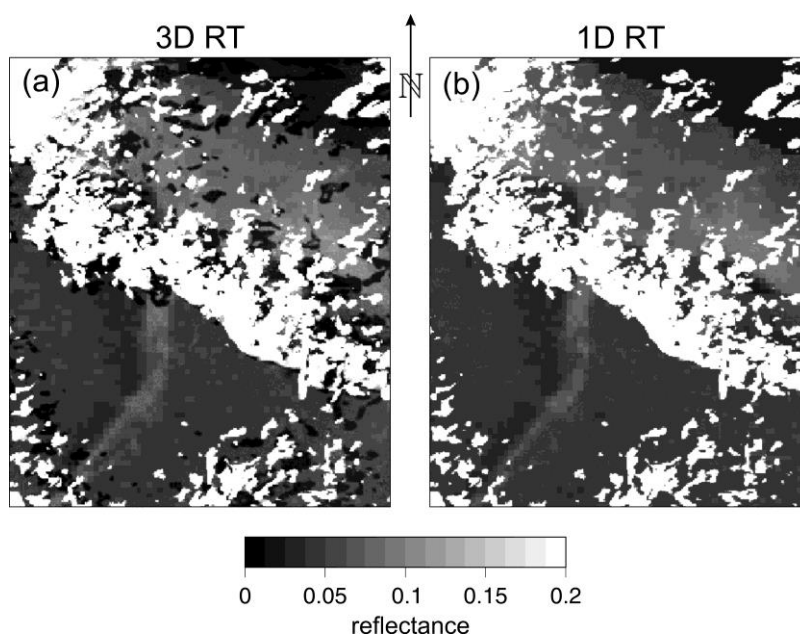
Lake breezes that arise from solar heating of land surfaces are simulated by a Numerical Weather Prediction (NWP) model using 250 m horizontal grid-spacing and either its regular 1D-ICA solar RT model (*control*) or the 1D parametrization (*experiment*). Idealized conditions are: i) flat surfaces; ii) lake with uniform temperature and optical properties; iii) uniform land-type with variable temperature, moisture, and albedo; iv) infinitely long linear coastlines running E-W; v) cyclic boundary conditions in the N-S direction with alternating strips of land and lake at 30 km and 70 km wide, respectively; vi) uniform and constant imposed synoptic winds; and vii) solar geometry for 43°N latitude on 8-July. Simulations started at 8h00 local time and lasted 8 hrs. Five member ensembles were produced for *control* and *experiment*. While the NWP model's response to forcings set-up by the 1D parametrization are clear and explicable, differences to surface meteorological variables are minor and would not impact weather forecasts. Owing to the high solar-sensitivity of these conditions, these results suggest that short-range NWP forecasts would gain little from having their efficient 1D RT models replaced by expensive 3D counterparts. This claim cannot, as yet, be extended to either seasonal forecasts or climate simulations.

23 1. Introduction

24 Sea breezes span a variety of sizes and strengths. At the top end are regional monsoons that usu-
25 ally last several months (e.g., the Indian Monsoon). At the bottom, localized lake breezes usually
26 last less than half a day. All such breezes are characterized by low-level flows of air from a cool
27 body of water that are drawn onto relatively warm adjacent land where temperatures are elevated
28 by solar radiative heating. Lake breezes frequent the coasts of North America's Great Lakes
29 during late-spring and summer. Shortly after mid-day, when surface solar irradiances maximize,
30 convective clouds form along the leading edge(s) of a lake breeze as moist low-level air from
31 over the lake ascends, cools, and condenses. In some cases, cells of intense convective precipita-
32 tion occur. Locally reduced surface temperatures due to cloud shadows, and possibly precipita-
33 tion, throttle-down lake breeze circulation thereby suppressing convection and cloud formation.
34 With fewer clouds, surface solar irradiance can increase near the weakened breeze-front, which
35 can reinvigorate it thus affecting a localized feedback process (cf. Gronemeier et al. 2017).

36 Numerical simulations of sea and lake breezes have been performed for many decades (e.g.,
37 Estoque 1961; Das 1980). When run with sufficiently small horizontal grid-spacings Δx , numer-
38 ical weather prediction (NWP) models simulate lake breezes well (Dehghan et al. 2018). NWP
39 models compute solar fluxes, however, with 1D solutions of the radiative transfer (RT) equation,
40 and so clouds cast shadows into nadir regardless of solar zenith angle θ_0 . Application of 1D RT
41 models to each column of a discretized domain is known as the Independent Column Approxima-
42 tion (ICA) (e.g., Barker and Davis 2005). As the 1D-ICA often yields accurate temporal-spatial
43 integrals of surface radiation budgets (e.g., Hogan et al. 2019), it is not surprising that the basic
44 features of monsoons can be simulate well. For localized, short-lived lake breezes, however, use
45 of the 1D-ICA might be problematic due to confinement of cloud shadow to nadir. Indeed, some

46 studies suggest that 3D RT effects can be important when simulating shallow cumulus clouds
 47 (Schumann et al. 2002; Wapler and Mayer 2008; Jakub and Mayer 2017). *Figure 1* illustrates this
 48 issue for Sun shining from the NW at $\theta_0 = 45^\circ$ and nadir-viewing (from satellite). In (a), 3D RT
 49 was performed, so shadows cast by clouds towards the SE are readily apparent. In (b), however,
 50 1D-ICA RT was performed, so all cloud shadows are hidden from view (perpetual opposition
 51 effect).



52
 53
 54 Figure 1. (a) Nadir-view of clouds (saturated white) above the NE coast of South America com-
 55 puted by a 3D solar RT model with the Sun coming in from the NW at $\theta_0 = 45^\circ$. (b) As in (a)
 56 except this was computed by a 1D RT model in ICA-mode.
 57

58 The purpose of this study is to explore whether the 1D-ICA's improper projection of cloud
 59 shadows impacts simulation of lake breezes for conditions resembling those near the Great Lakes
 60 of Southern Ontario. Experiments were performed with the Global Environmental Multi-scale
 61 (GEM) NWP model using first its regular ICA solar RT model, and then simple adjustments to
 62 ICA direct- and diffuse-beam surface irradiances. A full 3D solar RT model was not used for two

63 reasons: i) domain-average flux profiles computed by 1D-ICA models are often very good ap-
64 proximations of their 3D RT model counterparts (e.g., Hogan et al. 2019); and ii) for domains as
65 large as those used here, computer resources needed to perform simulations with a un-optimized
66 3D RT model would far exceed those used by the 1D-ICA; which already accounts for a signifi-
67 cant portion of the NWP model's run-time (per. comm., P. Vaillancourt 2018). Nevertheless,
68 these experiments should help establish the need to perform more extensive tests involving full
69 3D RT models.

70 The following section describes adjustments to ICA surface solar irradiances that aim to mim-
71 ic 3D RT values. These are assessed, in the third section, against 3D RT model results for a range
72 of partly cloudy model atmospheres. The fourth section provides a brief description of NWP
73 model used here for lake breeze experiments. This is followed by results and a conclusion.

74 **2. A simple transformation of ICA surface solar irradiances**

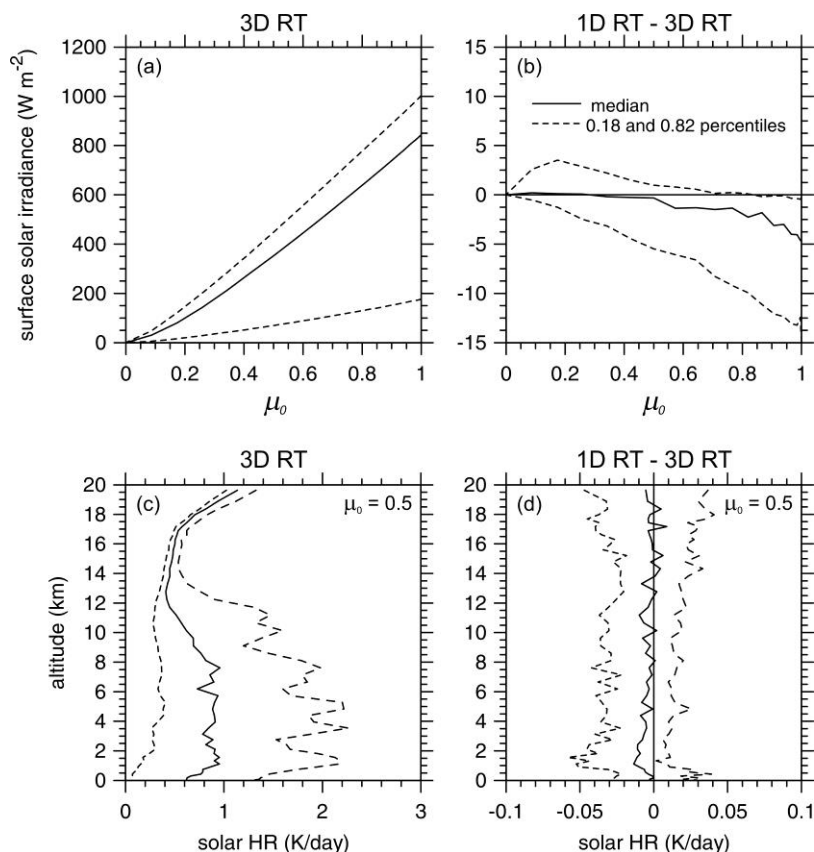
75 The underlying assumption going into this study was that poor estimation of 2D distributions of
76 solar heating over land by 1D-ICA RT models adversely impacts forecasts of lake breezes. To
77 facilitate tractable experiments aimed at helping establish whether NWP models should replace
78 their ubiquitous 1D solar RT models for more costly 3D solutions, simple adjustments to 1D-ICA
79 surface solar irradiances are proposed and explained in this section. Consider first, however, the
80 motivation behind the proposed adjustments.

81 **2.1. Parametrization motivation**

82 Results shown in this section, and in §3, were produced by applying a 3D Monte Carlo solar RT
83 model, with gaseous attenuation properties based on the RRTMG model (Iacono et al. 2008), to

84 sixty-five $(100 \text{ km})^2$ model-generated cloudy-sky domains above uniform ocean with horizontal
85 grid-spacing $\Delta x = 250 \text{ m}$. These domains, and the RT model, were also used by Hogan et al.
86 (2019) and were extracted from two large simulations: one from Greenland to Dominican Repub-
87 lic on 7 December 2014, and the other from Hawaii to Tonga on 24 June 2015. (Illingworth et al.
88 2015; Qu et al. 2023). Full 3D RT benchmarks used $\Delta x = 250 \text{ m}$, while their 1D-ICA counter-
89 parts used $\Delta x = 10^6 \text{ m}$, which affects the ICA via near-complete elimination of photon exchange
90 amongst columns.

91 While several studies have highlighted pathological differences in domain-average solar fluxes
92 estimated by 1D and 3D solar RT models for select cloud scenarios (e.g., Barker et al. 1999),
93 more holistic studies, that considered wide ranges of cloud forms, arrived at differences for sur-
94 face fluxes that are typically less than $\pm 20 \text{ W m}^{-2}$ (e.g., Ham et al. 2014; Barker et al. 2016;
95 Hogan et al. 2019; Gristey et al. 2020; Cole et al. 2023). These are echoed in *Figure 2* which
96 shows a repackaging of Hogan et al.'s (2019) results. *Figure 2a* and *Figure 2c* show median and
97 standard percentile values for domain-average global irradiance and solar heating rates, at cosine
98 of solar zenith angle $\mu_0 = \cos \theta_0 = 0.5$, where is θ_0 solar zenith angle, computed by the 3D RT
99 model for all 65 domains. Comparing these two plots to their counterparts on the right shows that
100 differences between 1D and 3D RT, at these scales, are much smaller than their median values.
101 For both surface irradiance and heating rates, 1D - 3D differences are usually less than 1%.

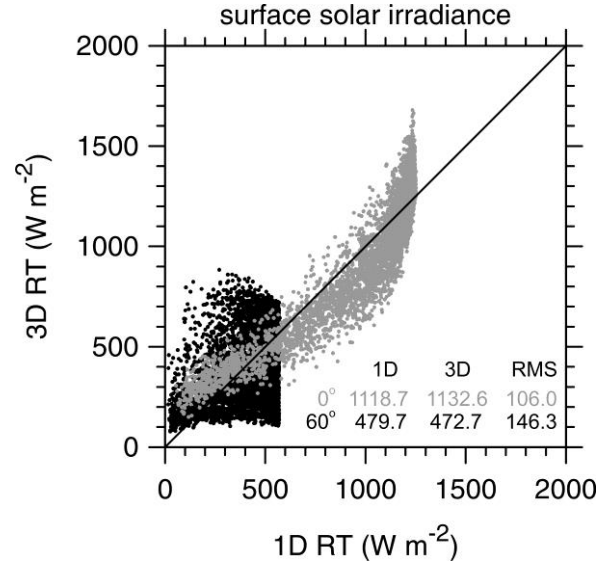


102

103 Figure 2. (a) Solid line is median of domain-average surface solar irradiance for 65 domains (see
 104 text) as a function of cosine of solar zenith angle μ_0 . Dashed lines are corresponding 0.18 and
 105 0.82 percentiles ($\sim 66\%$ of cases are between these lines; cf. standard deviation). (b) As in (a)
 106 except these are for differences between 1D and 3D RT results. (c) As in (a) except these are for
 107 HRs at $\mu_0 = 0.5$. (d) As in (b) except these are for heating rate differences.

108

109 Compare now values in **Figure 2** to those in **Figure 3** which shows, for a single typical do-
 110 main, 1D RT global irradiances against their 3D counterparts when averaged up to 1 km resolu-
 111 tion (which reduces Monte Carlo uncertainty by a factor of ~ 4). For $\theta_0 = 0^\circ$ (overhead Sun), 1D
 112 and 3D domain-averages differ by $\sim 1\%$, as in **Figure 2**, but locally they differ typically by $\sim 10\%$.
 113 This is due to horizontal diffusion of radiation scattered by clouds in the downwelling direction.
 114 At $\theta_0 = 60^\circ$, domain-averages still differ by only $\sim 1\%$ but locally they differ usually by $\sim 30\%$.
 115 This is due primarily to massively incorrect placement of direct-beam irradiance by 1D RT.



116

117 Figure 3. Each dot represents an ordered pair of surface solar irradiances for a 1 km square for a
 118 single domain as predicted by 1D RT and 3D RT. Grey and black dots are for $\theta_0 = 0^\circ$ and 60° ,
 119 respectively. Listed values are domain averages for both RT models with “RMS” standing for
 120 root mean-square difference between the two at 1 km resolution.

121

122 The null hypothesis arising from these results, and motivated development of the following
 123 parametrization, is: neglect by NWP models of 3D solar RT effects in the sub-mesoscale range
 124 outweigh corresponding errors at larger scales and does not affect weather forecasts.

125 2.2. Parametrization development

126 Define domain-average global surface irradiance predicted by a 1D solar RT model as

$$\begin{aligned}
 \langle G(\mu_0) \rangle &= \langle d(\mu_0) \rangle + \langle D(\mu_0) \rangle \\
 &= (1 - A_c) [\langle d_{clr}(\mu_0) \rangle + \langle D_{clr}(\mu_0) \rangle] + A_c [\langle d_{cld}(\mu_0) \rangle + \langle D_{cld}(\mu_0) \rangle],
 \end{aligned}
 \tag{1}$$

128 where $\langle d \rangle$ and $\langle D \rangle$ are domain-average diffuse- and direct-beam irradiances, A_c is vertically-
 129 projected total cloud fraction, $\langle d_{clr} \rangle$ and $\langle d_{cld} \rangle$ are mean diffuse-beam irradiances for the cloud-

130 less and cloudy portions of the domain, respectively, and $\langle D_{clr} \rangle$ and $\langle D_{cld} \rangle$ are their direct-beam
 131 counterparts. Correspondingly, results from a full 3D RT model can be cast as

$$\begin{aligned}
 \langle G'(\mu_0, \varphi_0) \rangle &= \langle d'(\mu_0, \varphi_0) \rangle + \langle D'(\mu_0, \varphi_0) \rangle \\
 &= [1 - A'_c(\mu_0, \varphi_0)] [\langle d'_{clr}(\mu_0, \varphi_0) \rangle + \langle D'_{clr}(\mu_0, \varphi_0) \rangle] \\
 &\quad + A'_c(\mu_0, \varphi_0) [\langle d'_{cld}(\mu_0, \varphi_0) \rangle + \langle D'_{cld}(\mu_0, \varphi_0) \rangle],
 \end{aligned} \tag{2}$$

133 where φ_0 is solar azimuth angle, and $A'_c(\mu_0, \varphi_0)$ is cloud fraction presented to the direct-beam, in
 134 which $A'_c(\mu_0 = 1) = A_c$. This is also the interpretation of the 1D ICA parametrization.

135 The first assumption in the 1D ICA parametrization is that 3D RT diffuse-beam fields can be
 136 approximated by simply smoothing their 1D RT compeers and constraining

$$\langle d(\mu_0) \rangle = \langle d'(\mu_0, \varphi_0) \rangle. \tag{3}$$

138 This allows (2) to be rewritten as

$$\begin{aligned}
 \langle G'(\mu_0, \varphi_0) \rangle &= \langle d(\mu_0) \rangle + [1 - A'_c(\mu_0, \varphi_0)] \langle D'_{clr}(\mu_0, \varphi_0) \rangle \\
 &\quad + A'_c(\mu_0, \varphi_0) \langle D'_{cld}(\mu_0, \varphi_0) \rangle.
 \end{aligned} \tag{4}$$

140 Admittedly, under some conditions, (3) can be extreme (e.g., Barker et al. 1999).

141 The second, and easily justifiable, assumption is that *mean* cloudless-sky direct-beam irradi-
 142 ances are equal for 1D and 3D RT. This, as in (3), implies that

$$\langle D_{clr}(\mu_0) \rangle = \langle D'_{clr}(\mu_0, \varphi_0) \rangle. \tag{5}$$

144 Next, assume further that domain-average global irradiances for the default and parametrized 1D
 145 RT models are equal. Thus, substituting

146
$$\langle G(\mu_0) \rangle = \langle G'(\mu_0, \varphi_0) \rangle \quad (6)$$

147 into (4) and using (5), mean cloudy-sky direct-beam irradiance for the parametrization is

148
$$\langle D'_{clد}(\mu_0, \varphi_0) \rangle = \frac{\langle G(\mu_0) \rangle - \langle d(\mu_0) \rangle - [1 - A'_c(\mu_0, \varphi_0)] \langle D_{clr}(\mu_0) \rangle}{A'_c(\mu_0, \varphi_0)}, \quad (7)$$

149 where everything on the RHS, save for $A'_c(\mu_0, \varphi_0)$, comes from the usual application of the 1D

150 RT model. The following two subsections describe estimation of $A'_c(\mu_0, \varphi_0)$ and definition of

151 spatial distributions of diffuse- and direct-beams for the parametrized 1D model.

152 2.3. Direct-beam irradiance and μ_0 -dependent cloud fraction

153 A ray-tracing algorithm, pared-down from Barker et al.'s (2003) Monte Carlo algorithm, is used

154 to approximate $A'_c(\mu_0, \varphi_0)$ and the location of parametrized direct-beam irradiance onto the

155 surface. From the centre of each surface grid-cell at (i, j) , a single ray is traced toward the Sun

156 and visible optical depth of cloud $\tau'_{clد}(i, j; \mu_0, \varphi_0)$ accumulated along it. This leads to location-

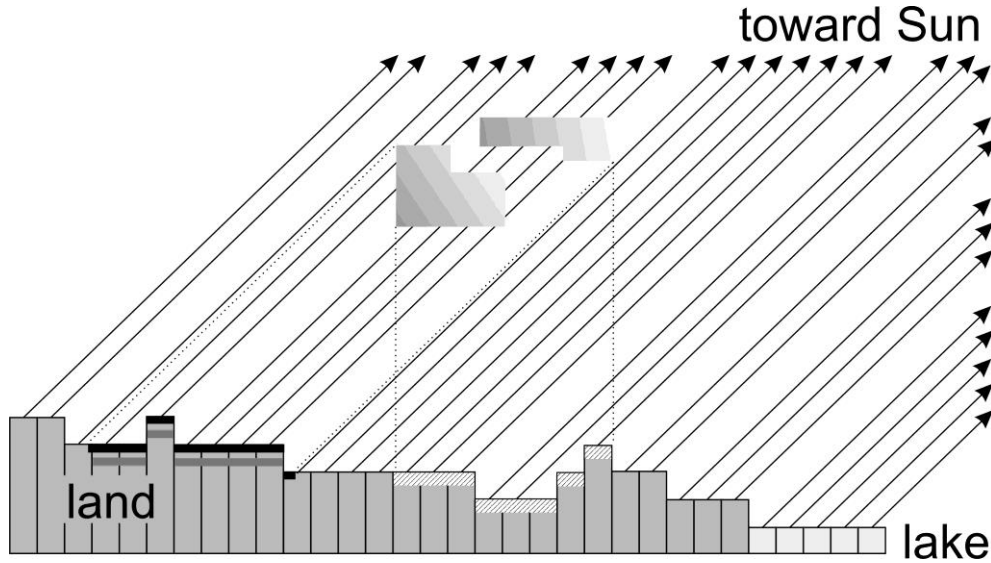
157 dependent direct-beam transmittances for clouds defined as

158
$$T'_{clد}(i, j; \mu_0, \varphi_0) = \begin{cases} \exp[-\tau'_{clد}(i, j; \mu_0, \varphi_0)] & ; \tau'_{clد} > \tau_{crit} \\ 1 & ; \tau'_{clد} \leq \tau_{crit}. \end{cases} \quad (8)$$

159 Hereinafter, $\tau_{crit} = 0.1$ was used. To ensure (6), direct-beam irradiance at a cloudy cell (i, j) for

160 the parametrized model is

161
$$D'_{clد}(i, j; \mu_0, \varphi_0) = \left[\frac{\langle D'_{clد}(\mu_0, \varphi_0) \rangle}{\langle T'_{clد}(i, j; \mu_0, \varphi_0) \rangle} \right] T'_{clد}(i, j; \mu_0, \varphi_0). \quad (9)$$



162

163 Figure 4. Schematic of the surface centre-point ray tracing algorithm used to estimate total cloud
 164 fraction presented to direct-beam A'_c , T'_{cl} (see (8)), and location of cloud shadows. Location of
 165 true cloud shadows are indicated in black while those for the 1D-ICA parametrization are dark
 166 grey. Cloud shadows for the true 1D-ICA are stippled bands.

167

168 For a domain of $N \times N$ columns, directional-dependent total cloud fraction presented to di-
 169 rect-beam is

$$170 \quad A'_c(\mu_0, \varphi_0) = \frac{1}{N^2} \sum_{i=1}^N \sum_{j=1}^N \Phi[\tau'_{cl}(i, j; \mu_0, \varphi_0)], \quad (10)$$

171 where

$$172 \quad \Phi[x] = \begin{cases} 0 & ; \quad x \leq \tau_{crit} \\ 1 & ; \quad x > \tau_{crit} \end{cases} \quad (11)$$

173 **Figure 4** shows a schematic of this procedure. From this simply illustration one can easily appre-

174 ciate that for small horizontal grid-spacings, errors in shadow locations and A'_c will be small.

175 2.4. Diffuse-beam irradiance

176 Denote 2D distributions of diffuse irradiance computed by 3D and 1D solutions of the RTE as
 177 $d'(i, j; \mu_0, \varphi_0)$ and $d(i, j; \mu_0)$, respectively. Since $d'(i, j; \mu_0, \varphi_0)$ consist of photons arriving from
 178 anywhere in the hemisphere, while $d(i, j; \mu_0)$ receives them only from directly above, a first-
 179 order approximation to $d'(i, j; \mu_0, \varphi_0)$ is to smooth $d(i, j; \mu_0)$. A very simplistic approach (cf.
 180 Marshak et al. 1995; Barker and Marshak 2001; Tjihuis et al. 2022) was taken here in which
 181 diffuse irradiance at (i, j) for the 1D-ICA parametrization (i.e., for 3D RT) is approximated as

$$182 \quad d'(i, j; \mu_0) = \frac{\sum_{m=-\ell}^{\ell} \gamma(m) d(i+m, j; \mu_0) + \sum_{m=-\ell}^{\ell} \gamma(m) d(i, j+m; \mu_0)}{2 \sum_{m=-\ell}^{\ell} \gamma(m)}, \quad (12)$$

183 where

$$184 \quad \gamma(k) = \cos^2 \left(\frac{k \cdot \Delta x}{h'_{cld}} \right), \quad (13)$$

185 ℓ corresponds to the number of cells removed from (i, j) where $\gamma(k) \leq \gamma_{crit}$, and h'_{cld} is an
 186 “effective” cloud altitude. Maximum weight is always given to the zenith with less weighty
 187 contributions coming from neighbouring columns along two perpendicular tracks that intersect
 188 (i, j) . For this study, γ_{crit} was set arbitrarily to 0.25. At this level of approximation, (12) is as
 189 justified as the more taxing use of all values of d in a circle of radius ℓ around (i, j) .

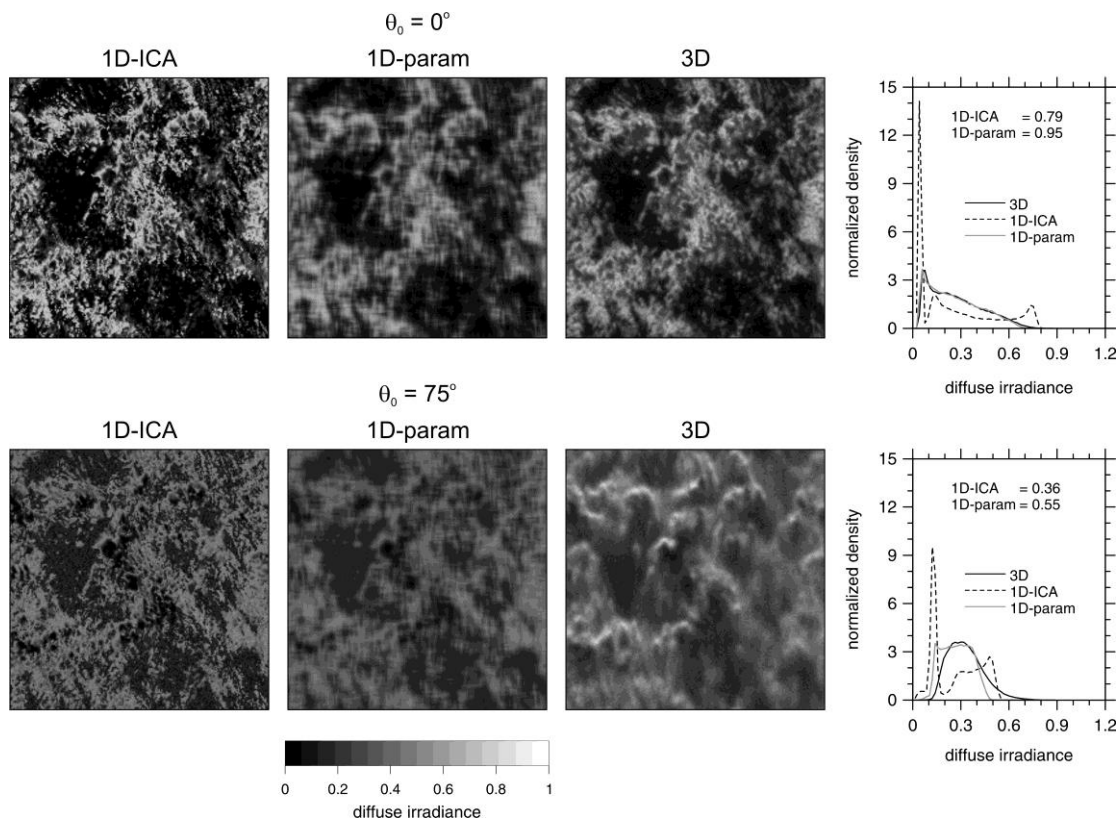
190 Letting $\langle w_{cld}(k) \rangle$ be the k^{th} layer’s domain-average cloud water content, and $\Delta z(k)$ be layer
 191 geometric thickness, h'_{cld} is defined as

$$192 \quad h'_{cld} \equiv z \left(\arg \max_k \left[\langle w_{cld}(k) \rangle \Delta z(k) \right] \right). \quad (14)$$

193 This approximation has almost no impact for near-cloudless skies and only minor impacts for
 194 most overcasts. While a better description of h'_{cld} would see it depend on (i, j) , its computation
 195 would not be justified for this level of approximation, and it could violate (3) and (6), which
 196 would upset experimental control. The final, and obvious, step is definition of local parametrized
 197 global irradiances being the addition of (12) to either $D'_{cld}(i, j)$ or $D_{clr}(i, j)$.

198 **3. Results I: Radiative transfer parametrization**

199 This section shows results for the RT parametrization presented in §2. All results in this section
 200 were produced by applying a 3D Monte Carlo solar RT model, with gaseous attenuation proper-
 201 ties based on the RRTMG model (Iacono et al. 2008), to sixty-five $(100 \text{ km})^2$ cloudy-sky do-
 202 mains, which were simulated by GEM, above a uniform ocean surface, with horizontal grid-
 203 spacing $\Delta x = 250$ m. These domains and the RT model (Cole et al. 2023) were also used in
 204 Hogan et al. (2019). Full 3D RT benchmarks used $\Delta x = 250$ m, while their 1D-ICA and para-
 205 metrization counterparts used $\Delta x = 10^6$ m, which affects the ICA via near-elimination of photon
 206 exchange between columns. Each simulation used 10^8 photons per domain. These scenes are
 207 subsets of two large domains that were simulated by GEM for the purpose of assessing Earth-
 208 CARE satellite retrieval algorithms (Illingworth et al. 2015; Qu et al. 2023): Greenland to Do-
 209 minican Republic on 7-12-2014; Hawaii to Tonga on 24-6-2015.



210

211 Figure 5. Upper row: Maps show surface diffuse irradiance, for overhead solar illumination
 212 ($\theta_0 = 0^\circ$), across a $(100 \text{ km})^2$ domain, from the Atlantic swath, predicted by the 1D-ICA, modi-
 213 fied 1D, and 3D RT models. Plot on the right shows distributions of values in the maps. Listed
 214 values are Pearson correlation coefficients between 1D RT and corresponding 3D values. Lower
 215 row: As in the upper except that $\theta_0 = 75^\circ$ with Sun coming in from 12 o'clock.

216

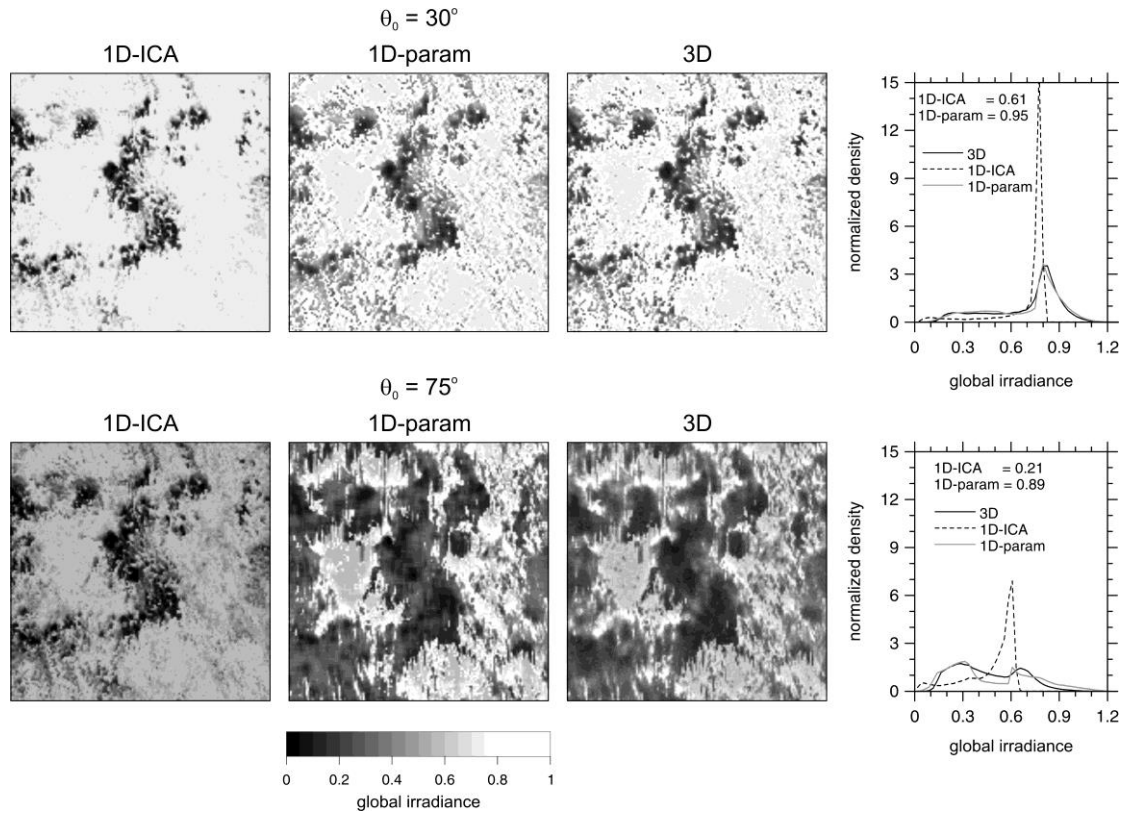
217 **Figure 5** shows maps of diffuse irradiance predicted by the 3D Monte Carlo, regular 1D-ICA,
 218 and parametrized ICA models for $\theta_0 = 0^\circ$ and $\theta_0 = 75^\circ$. At $\theta_0 = 0^\circ$, features in the 1D-ICA field
 219 are very “sharp”, while those in the 1D-param field are overly diffuse relative to the 3D. Maps for
 220 the 1D and 1D-param at $\theta_0 = 75^\circ$ look much the same as at 0° , whereas the 3D map is altered
 221 much due to cloud side-illumination and shadows. Nevertheless, Pearson correlation coefficients
 222 r between 1D model results and 3D values, as listed on the adjacent line plots, indicate that the
 223 smoothed 1D-ICA fields track the 3D results substantially better than the regular 1D-ICA.

224 Density functions in *Figure 5* indicate that 1D RT local irradiances are often too small or too
225 large; endless clear- or overcast-skies, respectively. The smoothing process described in §2.4,
226 however, yields distributions that agree extremely well with the 3D model's. Note that the rela-
227 tively long tail for 3D RT at $\theta_0 = 75^\circ$ stems from enhanced interception of direct-beam radiation
228 by clouds and the ensuing concentrating of radiation scattered downward onto nearby ground.

229 *Figure 6* is like *Figure 5* except it shows global surface irradiances. Clearly, maps of para-
230 metrized 1D results resemble the 3D maps much better than do the regular 1D-ICA. This is re-
231 flected in r which tend to increase by 0.4 to 0.6 thanks to the vast improvement in positioning of
232 direct-beam irradiance across the domain. As with the diffuse fields, frequency distributions for
233 1D-param resemble closely those for the 3D RT, while the regular 1D-ICA's distributions are
234 remarkably narrow (cf. Barker et al. 2017).

235 *Figure 7* shows mean and standard deviations of r for fields of surface irradiances for the 1D-
236 ICA and 1D-param models relative to their 3D RT counterparts as functions of μ_0 for all 65 (100
237 km)² domains. For diffuse irradiance, parametrized 1D values of r are about 0.2 larger than those
238 for the 1D-ICA, regardless of θ_0 (cf. *Figure 5*). More importantly, especially for non-overcast
239 domains, direct-beam for 1D-param are almost perfectly correlated with the 3D benchmarks, and
240 this translates into values of r between 1D-param and 3D fields rarely being less than 0.8 and
241 always being much larger than those for the regular 1D-ICA model.

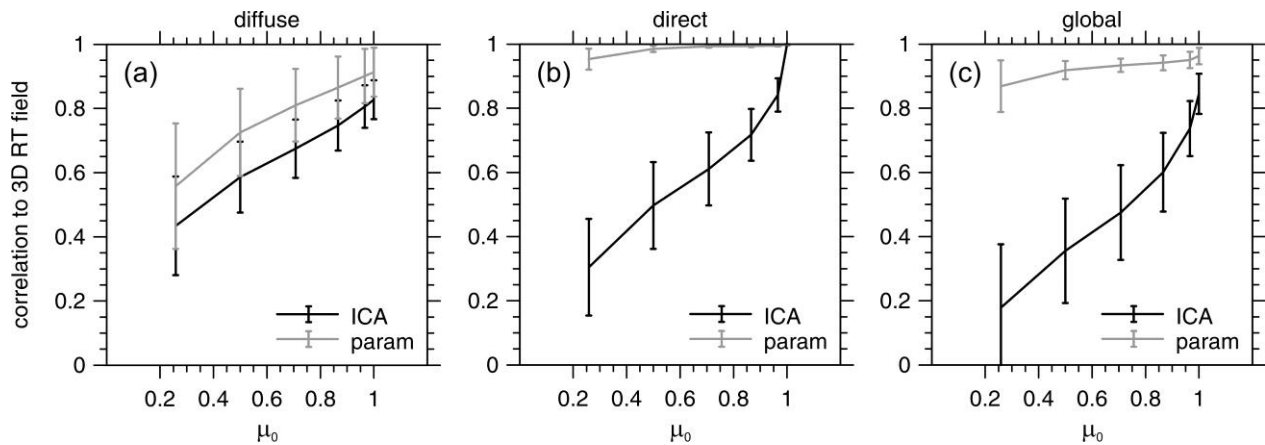
242 In summary, biases incurred by 1D-ICA RT models most in need of addressing are those at
243 small-scales for partly cloudy skies. To this end, the simple adjustment to 1D-ICA surface irradi-
244 ances should suffice in exploring the impact of neglecting, expensive to obtain, 3D RT effects.
245 The following sections do this for solar-driven lake breeze conditions.



246

247 Figure 6. As in *Figure 5* but these are for surface global irradiances at $\theta_0 = 30^\circ$ and $\theta_0 = 75^\circ$.

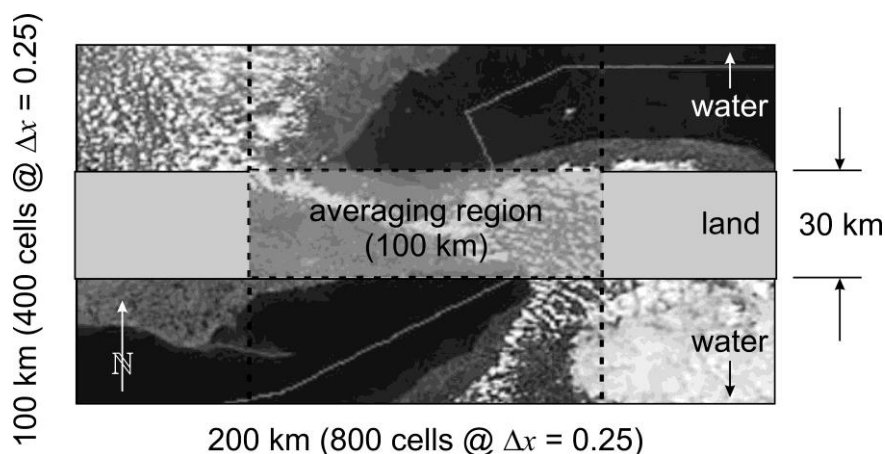
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249

250 Figure 7. Pearson correlation coefficients r between 2D fields of diffuse, direct, and global sur-
 251 face irradiances predicted by the ICA and the parametrized ICA RT models relative to their 3D
 252 RT counterparts, as functions of μ_0 . Solid lines indicate mean values for the 65 scenes consid-
 253 ered in this portion of this study. Corresponding standard deviations are indicated by the bars.

254



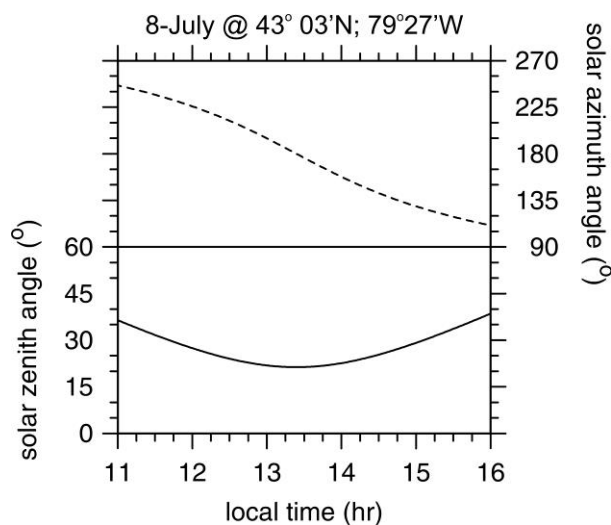
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256 Figure 8. Schematic showing GEM's inner-domain (200 km x 100 km) with a 30 km-wide strip
 257 of land bordered by water. It is superimposed on a satellite image of the Niagara region between
 258 Lakes Ontario and Erie. For analyses, only the central 100 km x 30 km portion is considered.

259

260 4. NWP model simulation of lake breezes

261 For this study, ECCC's Global Environmental Multiscale (GEM) NWP model was used (Côté et
 262 al. 1998; Girard et al. 2014). For details, including radiation and cloud-precipitation microphys-
 263 ics, see McTaggart-Cowan et al. (2019). GEM is usually run with four one-way nested domains
 264 (Milbrandt et al. 2016; Bélair et al. 2017). Experiments reported here, however, were performed
 265 with an idealized version using $\Delta x = 0.25$ km and 57 vertical layers with inner-domain measur-
 266 ing 200 km east-west by 100 km north-south. There was a continuous and homogenous strip of
 267 land running east-west that was 30 km north-south. Uniform water bordered it, and the entire
 268 domain was set in a larger water surface domain with $\Delta x = 1$ km. As **Figure 8** shows, this con-
 269 figuration resembles the isthmus between Lakes Ontario and Erie where lake breezes are com-
 270 mon. Domains were cyclic and analyses were restricted to the central 100 km of land. At the end
 271 of a time step, the 1D-ICA RT model's direct- and diffuse surface irradiance fields were either: i)
 272 used as usual (i.e., *control*); or ii) adjusted according to that described in §2 (*experiment*).



273

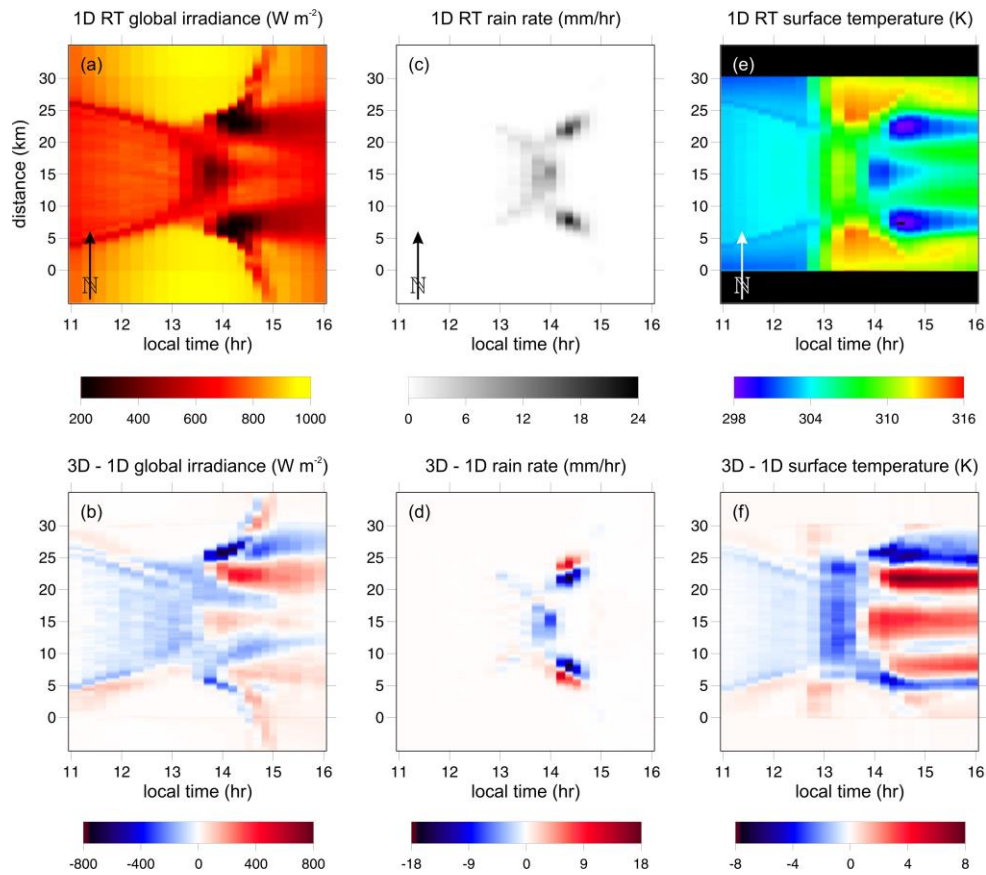
274 Figure 9. Solar zenith and azimuth angles as functions of local time for the date and location
 275 listed in the title. Solar noon occurs at about 13h20 (daylight saving time) when $\varphi_0 = 180^\circ$.

276

277 All simulations started at 12h00 UTC using initial conditions from archived analysis for 8-
 278 July-2020. The centre of the domain corresponds to $(43^\circ 03' N; 79^\circ 27' W)$, and **Figure 9** shows
 279 solar zenith and azimuth angles. Synoptic winds were imposed from the SW, NW, or W direc-
 280 tions and fixed during simulations. For the W-wind, the westerly component was from analysis
 281 data and reduced by 50% in order to establish a lake breeze. The southerly component was set to
 282 0. For SW and NW cases, the north/south component was prescribed at 25% of the westerly
 283 component. A fourth scenario had no synoptic wind. Most of the results shown in the following
 284 section pertain to the “no synoptic wind” and “SW” experiments because they capture the essen-
 285 tial features and these conditions are common when lake breezes prevail near the Great Lakes.

286 For each wind direction five member ensembles were performed for both *control* and *experi-*
 287 *ment*. Each ensemble member had initial values of land surface temperatures, from analysis,
 288 perturbed by random Gaussian noise with standard deviation 1°C . Water surface temperatures
 289 were fixed at 24°C .

290



291

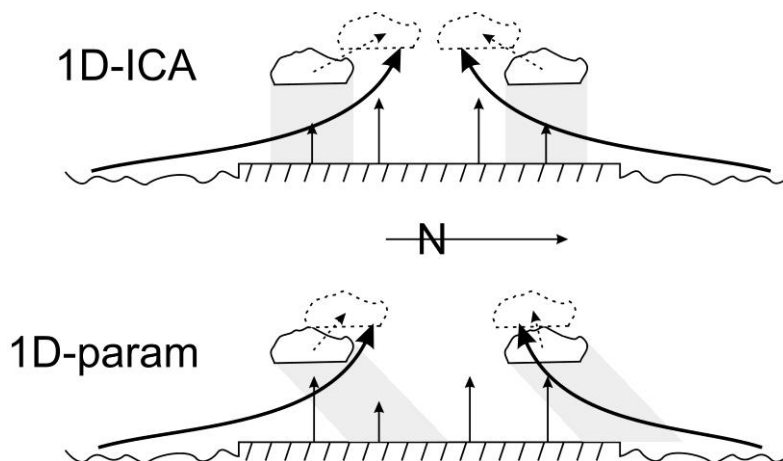
292 Figure 10. (a) Ensemble-average time series of global surface irradiance averaged longitudinally
 293 across the “averaging region”, as indicated in *Figure 8*, for simulations using regular 1D RT for
 294 “no synoptic wind” conditions. (b) Difference between (a) and corresponding values using para-
 295 metrized 1D RT (referred to here as 3D RT). (c) and (d) are as in (a) and (b) except these are for
 296 surface rain rate. (e) and (f) are as in (a) and (b) except these are for surface temperature.
 297

298 5. Results

299 Consider first the case of ‘no synoptic wind’. For 1D RT, clouds cast shadows directly beneath
 300 themselves and so lake breeze conditions should be very close to symmetric about the W-E centre
 301 of the isthmus. This is confirmed in *Figure 10a* which shows a Hovmöller-style plot of global
 302 surface irradiance averaged W-E across the domain as a function local time. The plot begins

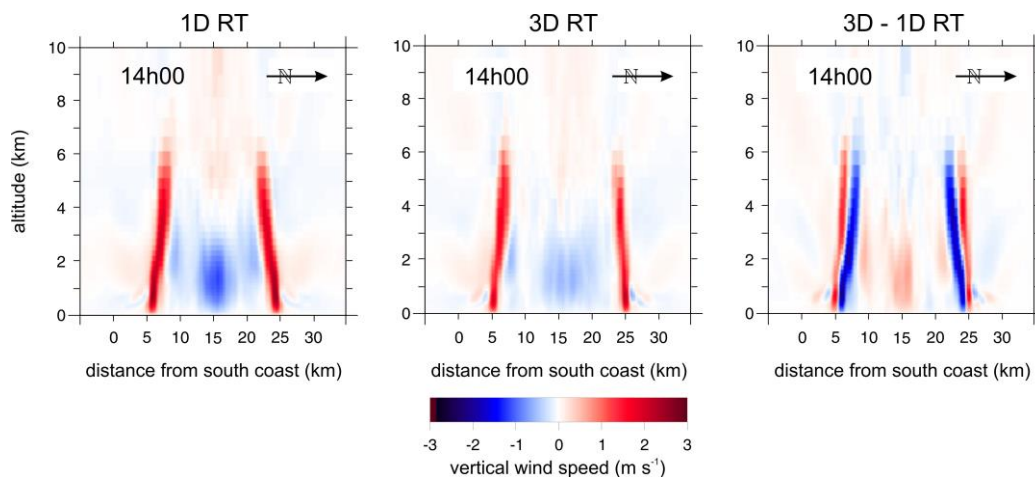
303 shortly after scattered shallow cumulus form inland of the lake breeze fronts; both of which are,
304 at this time, ~4 km from the coasts. Conditions are symmetric throughout the simulation with
305 irradiance minima occurring shortly after 14h00 along the breeze fronts, which have migrated to
306 ~7 km inland. This coincides with maximum surface rain rate (*Figure 10c*) which, coupled with
307 diminished surface solar heating, cools the surface (*Figure 10e*) almost shutting down the lake
308 breezes. With areas of rain dying and moving coastward, with clouds extending out over water
309 and lessening inland, mid-afternoon surface temperatures across most of the isthmus rise despite
310 the Sun going down (*Figure 9*).

311 The lower row of panels in *Figure 10* show differences between the top row's 1D RT results
312 and results for the simulations using parametrized 1D RT, which are referred to as "3D RT" in
313 the plot titles only because they are meant to approximate 3D RT. The mild asymmetry between
314 the north and south sides of the isthmus, as seen in *Figure 10b*, is set-up by the shifting of cloud
315 shadows in the direction of the south coast's breeze and counter to the north coast's breeze. A
316 schematic of this is shown in *Figure 11* which shows why clouds, and ultimately rain, that form
317 along the breeze fronts shift coastward for approximate 3D RT (cf. Gronemeier et al. 2017). This
318 is confirmed by *Figure 12* and *Figure 13* which show that, near peak-breeze strength at 14h00,
319 vertical wind speeds are generally weaker and cloud densities less for 3D RT. While these differ-
320 ences were the largest seen here, they would be of little concern to weather forecasters.

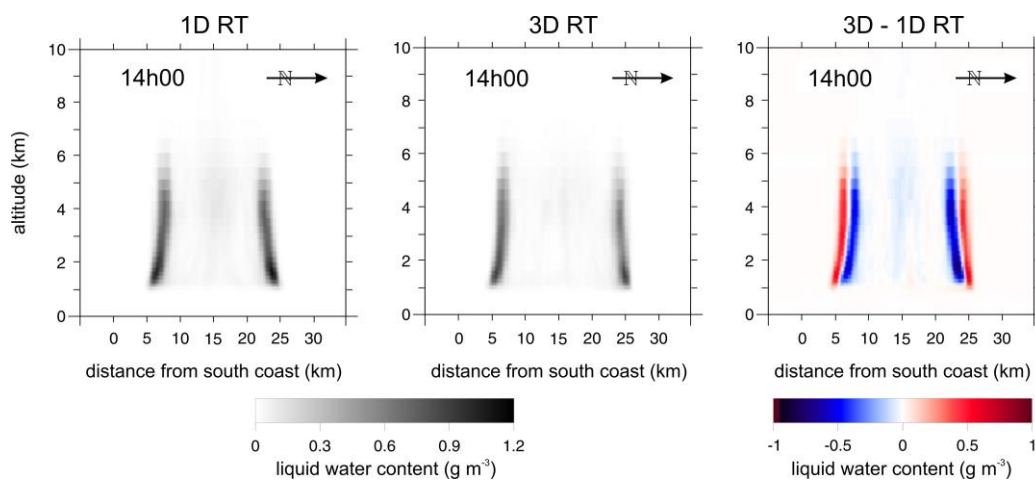


321
 322 Figure 11. Schematic showing clouds forming along lake breeze fronts for both RT treatments for
 323 “no synoptic wind” conditions. For the 1D RT model, clouds that form along the breeze front
 324 cast shadows directly beneath themselves. They are drawn in the direction of the breezes due to
 325 surface heating just inland of them. For the south shore’s breeze, this drawing inland is curtailed
 326 whilst heating remains to take place directly beneath clouds (when μ_0 is not too large). For the
 327 north breeze, heating beneath cloud also occurs but the breeze can be expected to be diminished
 328 due to cloud shadows cast further towards the coast.
 329

330



331
 332 Figure 12. Longitudinal average cross-sections of vertical wind speed averaged across the “aver-
 333 aging region” shown in **Figure 8** for simulations using regular 1D RT, parametrized (3D) RT,
 334 and their differences at 14h00 local time. These are for “no synoptic wind” conditions.
 335

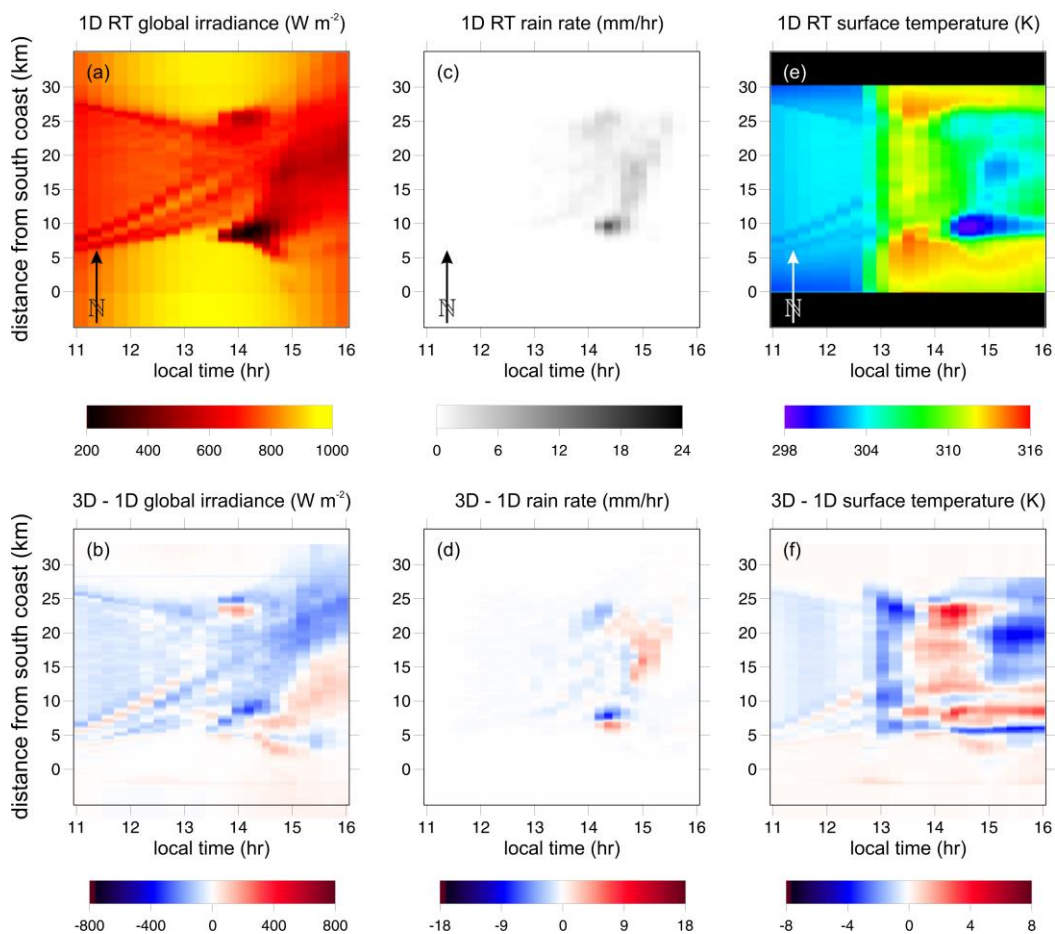


336

337

338

Figure 13. As in *Figure 12* except these show cloud liquid water contents.



339

340

Figure 14. As in *Figure 10* but these are for SW synoptic winds.

341 Although lake breezes along the Great Lakes are common when synoptic winds are weak dur-
342 ing high pressure conditions, they also occur during fair-weather conditions with south-westerly
343 winds. *Figure 14* is a repeat of *Figure 10* for SW synoptic winds. It is immediately clear that for
344 1D-ICA RT the SW flow has upset the symmetry relative to no synoptic wind; especially the
345 north shore breeze which is so diminished, fighting against the synoptic flow, that only a weak
346 initial band of rain develops along its front. Differences in surface global irradiance, rain rate, and
347 temperature affected by use of the 1D parametrization are also weakened relative to “no synoptic
348 wind”. What does develop, however, is enhanced rain, via enhanced convection (e.g., Veerman et
349 al. 2020), near the centre of the isthmus at ~15h00 with notable subsequent surface cooling due to
350 the combined effects of surface moistening and reduced irradiance, and north-easterly migration
351 of rain thanks to synoptic wind. But as with no synoptic winds, differences between 1D-ICA and
352 the pseudo-3D RT parametrization are minor and unlikely to alter short-term weather forecasts.

353 *Table 1* lists domain-averages and standard deviations of surface global irradiance, tempera-
354 ture, and rain rate when 1D and 3D RT simulations are integrated across all “averaging domains”
355 (see *Figure 8*) between 11h00 and 16h00 local time, for all synoptic wind experiments. In all
356 cases, 3D RT leads to 2% - 6% less surface irradiance due mostly to thicker, albeit slightly fewer,
357 clouds. The impact of thicker but few clouds is also manifest in enhanced standard deviations of
358 surface irradiance across the domain; direct-beam variations throttling global irradiances between
359 very large and small values (see *Figure 6*). While surface temperatures for 1D and 3D RT differ
360 negligibly, fractional areas with rain decrease slightly when 3D RT is used, with rain intensity
361 both increasing and decreasing, though by no more than $\pm 6\%$.

362

363 **Table 1. Domain-averages, domain-standard deviations, and fractional areas averaged over all**
 364 **ensemble members from 11h00 to 16h00 local time.**

	global irradiance (W m ⁻²)		surface temperature (K)		surface rain rate > 0 (mm hr ⁻¹)	
	<i>no synoptic wind</i>					
	1D	3D	1D	3D	1D	3D
domain-average	714.1	698.8	306.2	306.2	2.98	2.84
domain-stdev	255.1	305.2	4.8	4.7	9.34	9.02
area	-	-	-	-	0.21	0.17
	W					
	1D	3D	1D	3D	1D	3D
domain-average	748.8	701.8	306.7	306.4	2.91	2.81
domain-stdev	227.1	289.1	4.7	4.7	8.86	8.71
area	-	-	-	-	0.17	0.17
	NW					
	1D	3D	1D	3D	1D	3D
domain-average	747.4	706.3	306.7	306.5	2.76	2.83
domain-stdev	227.4	285.9	4.8	4.7	8.50	8.79
area	-	-	-	-	0.18	0.17
	SW					
	1D	3D	1D	3D	1D	3D
domain-average	745.9	705.7	306.8	306.6	2.90	3.06
domain-stdev	232.7	291.6	4.8	4.8	8.72	9.30
area	-	-	-	-	0.18	0.17

365

366 6. Summary and discussion

367 An algorithm was presented that makes simple, and numerically efficient, adjustments to spatial
 368 distributions of diffuse- and direct-beam surface solar irradiances as predicted by the standard
 369 1D-ICA RT framework, thus bringing them into better alignment with results from 3D radiative

370 transfer (RT) models. The modified-ICA is best applied in well-resolved NWP or cloud-resolving
371 models. The motivation for this was to help NWP modellers decide if it is necessary to replace
372 the ubiquitous 1D solar RT model, with a 3D counterpart, before expending much effort on im-
373 plementing computationally expensive 3D RT models in NWP models.

374 A smoothing filter is applied to the 2D distribution of surface diffuse irradiances simulated by
375 a 1D-ICA model to approximate lateral diffusion of radiation below clouds. The breadth of the
376 filter increases with cloud altitude, beginning with no smoothing for surface fog. Using high-
377 resolution cloudy-sky data (cf. Barker and Davies 1989; Hogan et al. 2019) it was shown that this
378 simple alteration typically increases the Pearson correlation coefficient between 1D-ICA and
379 truly 3D RT modelled diffuse-beam irradiances by 0.1 to 0.2. Direct-beam irradiances for the
380 modified-ICA are set by tracing a single ray from the centre of each surface cell toward the Sun,
381 accumulating cloud optical depth along them, and applying Beer's law. This positions cloud
382 shadows accurately and typically increases the Pearson correlation coefficient between 1D-ICA
383 and truly 3D RT direct-beam irradiances by 0.65 at $\theta_0 = 75^\circ$ and 0.5 at $\theta_0 = 60^\circ$; correctly, no
384 improvement occurs at $\theta_0 \approx 0^\circ$ as 1D RT cloud shadows are already positioned properly.

385 While the modified-ICA and standard 1D-ICA have equal domain-average diffuse irradiances,
386 which, strictly speaking, is incorrect (cf. Welch and Wielicki 1985), the modified-ICA's direct
387 irradiances are normalized by forcing its domain-average *global* irradiance to equal the standard
388 1D-ICA's. This neglect of differences between 1D-ICA and 3D mean surface irradiances should
389 be minor, for as several studies have shown (e.g., Pincus et al. 1999; Di Giuseppe and Tompkins
390 2003; O'Hirok and Gautier 2005; Ham et al. 2014; Barker et al. 2015, 2016; Hogan et al. 2019;
391 Gristey et al. 2020), these differences are typically less than 10 W m^{-2} (or $\sim 2\%$). Moreover, at-
392 mospheric heating rates for the modified-ICA are identical to those produced by the standard 1D-

393 ICA, which differ only slightly from their true 3D counterparts (see *Figure 2* and Hogan et al.
394 2019).

395 To gauge the importance of taking proper account of cloud shadows, the GEM NWP model
396 used the standard- and modified-ICAs to simulate lake breezes. Lake breezes were focussed on
397 because they are set-up when land surfaces, adjacent to cool lakes, are warmed by solar irradi-
398 ance. They often lead to lines of cumulus clouds and localized convective precipitation; both of
399 which feedback on the circulation that spawned them (e.g., Jakub and Mayer 2017). Since GEM,
400 like all NWP models, employs 1D-ICA models whose clouds cast shadows at nadir, regardless
401 of θ_0 , the standard- and modified-ICAs enabled testing of whether NWP model simulations of
402 lake breezes are sensitive to location of shadows cast by clouds on land surfaces.

403 In summary, while GEM's simulations of lake breeze-dependent cloud and surface meteoro-
404 logical conditions respond in clear and explainable ways when its standard 1D-ICA solar RT
405 model is replaced by its modified version, which captures important features of 3D RT, the im-
406 pacts are unlikely to alter local weather forecasts. Given the strong dependence of lake breezes on
407 local solar heating, it is hard to imagine other short-term weather scenarios being significantly
408 more sensitive to 3D solar RT effects (e.g., for deep convection, differences in radiative heating
409 due to 1D and 3D RT models are likely to be greatly overshadowed by other processes). Hence,
410 there is, as yet, no compelling reason for NWP models, which are used for short-range weather
411 forecasting, to abandon their efficient 1D-ICA RT models. It is too early to tell, however, if this
412 statement applies to seasonal weather and long-term climate predictions, where *cumulative effects*
413 of 1D-ICA biases might be important (see Hogan and Bozzo 2018). In the meantime, NWP model-
414 lers might be advised to consider employment of methods that aim to improve numerical effi-
415 ciency of 1D-ICA RT (e.g., Barker and Li 2019; Barker et al. 2020).

416 **Acknowledgements**

417 In the short-term, data used for this study can be obtained from the corresponding author. By the
418 time of publication, ECCC will have an official data repository from which data used here will be
419 downloadable.

420

421 **References**

- 422 Barker, H. W. and J. A. Davies, 1989: Multiple reflections across a linear discontinuity in surface albedo.
423 *Int. J. Clim.*, **9**, 203-214.
- 424 Barker, H. W., G. L. Stephens, and Q. Fu, 1999: The sensitivity of domain-averaged solar fluxes to as-
425 sumptions about cloud geometry. *Q. J. R. Meteorol. Soc.*, **125**, 2127-2152.
- 426 Barker, H. W. and A. Marshak, 2001: Inferring optical depth of broken clouds above green vegetation
427 using surface solar radiometric measurements. *J. Atmos. Sci.*, **58**, 2989-3006.
- 428 Barker, H. W., R. K. Goldstein, and D. E. Stevens, 2003: Monte Carlo simulation of solar reflectances for
429 cloudy atmospheres. *J. Atmos. Sci.*, **60**, 1881-1894.
- 430 Barker, H. W. and A. Davis, 2005: Approximation Methods in Atmospheric 3D Radiative Transfer, Part
431 2: Unresolved Variability and Climate. pg. 343-384 in *Three-dimensional Cloud Structure and Ra-*
432 *diative Transfer*. Eds. A. Marshak and A., Davis. Springer-Verlag, 686pp.
- 433 Barker, H. W., S. Kato, and T. Wehr, 2012: Computation of Solar Radiative Fluxes by 1D and 3D Meth-
434 ods using Cloudy Atmospheres Inferred from A-train Satellite Data. *Surveys in Geophys.* **33**, 657-676.
- 435 Barker, H. W., J. N. S. Cole, J. Li, B. Yi, and P. Yang, 2015: Estimation of Errors for Two-Stream Ap-
436 proximations of the Solar Radiative Transfer Equation for Cloudy Atmospheres. *J. Atmos. Sci.*, **72**,
437 4053-4074.
- 438 Barker, H. W., J. N. S. Cole, J. Li, and K. von Salzen, 2016: A Parametrization of 3D Subgrid-scale
439 Clouds for Conventional GCMs: Assessment using A-Train Satellite Data and Solar Radiative Transfer
440 Characteristics. *J. Adv. Model. Earth Syst.* DOI: 10.1002/2015MS000601.
- 441 Barker, H. W., Z. Qu, S. Belair, S. Leroyer, J. A. Milbrandt, P. A. Vaillancourt, 2017: Scaling Properties
442 of Observed and Simulated Satellite Visible Radiances. *J. Geophys. Res.*, **122**, 9413– 9429,
443 doi:10.1002/2017JD027146.
- 444 Barker, H. W. and J. Li, 2019: Accelerating radiative transfer calculations for high-resolution atmospheric
445 models. *Q. J. R. Meteorol. Soc.*, **145**, 2046–2069.
- 446 Barker, H. W., Z. Qu, V. Dhanraj, and J. N. S. Cole, 2020: Partial validation of a lossy compression
447 approach to computing radiative transfer in cloud system-resolving models. *Q. J. R. Meteorol. Soc.*,
448 **146**, 1–20. <https://doi.org/10.1002/qj.3922>
- 449 Cole, J. N. S., H. W. Barker, Z., Qu, N., Villefranque, and M. W. Shephard, 2023 : Broadband Radiative
450 Quantities for the EarthCARE Mission: The ACM-COM and ACM-RT Products. *Atmospheric Meas-*
451 *urement Techniques*, **16**, 4271–4288.
- 452 Côté, J., S. Gravel, A. Méthot, A. Patoine, M. Roch, and A. Staniforth, 1998: The operational CMC–MRD
453 Global Environmental Multiscale (GEM) model. Part I: Design considerations and formulation. *Mon.*
454 *Wea. Rev.*, **126**, 1373–1395.

- 455 Das, P. K., 1980: Numerical simulation of the Indian monsoon. *Proc. Indian Academy of Sci. - Earth and*
456 *Planet. Sci.*, **89**, 159-177.
- 457 Dehghan, A., Z. Mariani, S. Leroyer, P. Joe, D. Sills, S. Bélair, 2018: Evaluation of Modeled Lake Breez-
458 es Using an Enhanced Observational Network in Southern Ontario: Case Studies. *J. of Meteorol.*
459 *and Climatol.*, **57**, 1511–1534.
- 460 Di Giuseppe, F. and A. M. Tompkins, 2003: Effect of Spatial Organization on Solar Radiative Transfer in
461 Three-Dimensional Idealized Stratocumulus Cloud Fields. *J. Atmos. Sci.*, **60**, 1774-1794.
- 462 Estoque, M. A., 1961: A theoretical investigation of the sea breeze. *Q. J. R. Meteorol. Soc.*, **87**, 136-146.
- 463 Girard, C., M. Desgagné, R. McTaggart-Cowan, J. Côté, M. Charron, S. Gravel, V. Lee, A. Patoine, A.
464 Qaddouri, M. Roch, L. Spacek, M. Tanguay, P. A. Vaillancourt, and A. Zadra, 2014: Staggered
465 Vertical Discretization of the Canadian Environmental Multiscale (GEM) model using a coordinate
466 of the log-hydrostatic-pressure type. *Mon. Weather Rev.*, **142**, 1183-1196. doi.org/10.1175/MWR-
467 D-13-00255.1
- 468 Gristey, J. J., G. Feingold, I. B. Glenn, K. S. Schmidt, and H. Chen, 2020: Surface Solar Irradiance in
469 Continental Shallow Cumulus Fields: Observations and Large-Eddy Simulation. *J. Atmos. Sci.*, **77**,
470 1065–1080.
- 471 Gronemeier, T., F. Kanani-Sühring, and S. Raasch, 2017: Do Shallow Cumulus Clouds have the Potential
472 to Trigger Secondary Circulations Via Shading? *Boundary-Layer Meteo.*, **162**, 143–169.
- 473 Ham, S.-H., S. Kato, H. W. Barker, F. G. Rose, and S. Sun-Mack, 2015: Improving the modeling of
474 shortwave radiation through the use of a 3D scene construction algorithm. *Q. J. R. Meteorol. Soc.*,
475 **141**, 1870–1883. DOI: 10.1002/qj.2491.
- 476 Hogan, R. J. and A. Bozzo, 2018: A flexible and efficient radiationscheme for the ECMWF model. *J. Adv.*
477 *Model. Earth Sys.*, **10**, 1990-2008. <https://doi.org/10.1029/2018MS001364>
- 478 Hogan, R. J., M. D. Fielding, H. W. Barker, N. Villefranque, and S. A. K. Schafer, 2019: Entrapment: An
479 important mechanism to explain the shortwave 3D radiative effect of clouds. *J. Atmos. Sci.*, **76**,
480 2123-2141.
- 481 Iacono, M. J., J. S. Delamere, E. J. Mlawer, M. W. Shephard, S. A. Clough, and W. D. Collins, 2008:
482 Radiative forcing by long-lived greenhouse gases: Calculations with the AER radiative transfer
483 models, *J. Geophys. Res.*, **113**, D13103, doi:10.1029/2008JD009944.
- 484 Illingworth, A., + others, 2015: THE EARTHCARE SATELLITE: The next step forward in global meas-
485 urements of clouds, aerosols, precipitation and radiation. *Bull. Am. Met. Soc.*, **96**, 1311-1332.
486 DOI:10.1175/BAMS-D-12-00227.1.
- 487 Jakub, F. and B. Mayer, 2017: The role of 1-D and 3-D radiative heating in the organization of shallow
488 cumulus convection and the formation of cloud streets. *Atmos. Chem. Phys.*, **17**, 13317–13327.
489 <https://doi.org/10.5194/acp-17-13317-2017>
- 490 Marshak, A., A. B. Davis, W. J. Wiscombe, and R. F. Cahalan, 1995: Radiative smoothing in fractal
491 clouds. *J. Geophys. Res.*, **100**, 26 247–26 261.

- 492 McTaggart-Cowan, R., Vaillancourt, P. A., Zadra, A., Chamberland, S., Charron, M., Corvec, S., Mil-
493 brandt, J. A., Paquin-Ricard, D., Patoine, A., Roch, M., Separovic, L., and Yang, J., 2019: Moderni-
494 zation of atmospheric physics parameterization in Canadian NWP. *J. Adv. Model. Earth Syst.*, **11**,
495 3593–3635, <https://doi.org/10.1029/2019MS001781>.
- 496 Milbrandt, J. A., Bélair, S., Faucher, M., Vallée, M., Carrera, M. L., and Glazer, A., 2016: The pan-
497 Canadian high resolution (2.5km) deterministic prediction system, *Weather Forecast.*, **31**, 1791–
498 1816, <https://doi.org/10.1175/WAF-D-16-0035.1>.
- 499 O’Hirok, W. and C. Gautier, 2005: The impact of model resolution on differences between independent
500 column approximation and Monte Carlo estimates of shortwave surface irradiance and atmospheric
501 heating rate., *J. Atmos. Sci.*, **62**, 2939–2951, <https://doi.org/10.1175/JAS3519.1>.
- 502 Pincus, R., S. A. McFarlane, and S. A. Klein, 1999: Albedo bias and the horizontal variability of clouds in
503 subtropical marine boundary layers: Observations from ships and satellites. *J. Geophys. Res.*, **104**,
504 6183-6191.
- 505 Qu, Z., D. P. Donovan, H. W., Barker, J. N. S., Cole, and M. Shephard, 2023: Numerical model genera-
506 tion of test frames for pre-launch studies of EarthCARE’s retrieval algorithms and data manage-
507 ment system. *Atmospheric Measurement Techniques*, **16**, 2319-2331.
- 508 Schumann, U. A. Dörnbrack, and B. Mayer, 2002: Cloud-shadow effects on the structure of the convec-
509 tive boundary layer. *Meteorologische Zeitschrift*, **11**, 285-294.
- 510 Tjihuis, M., B. van Stratum, and C. C. van Heerwaarden, 2022: An efficient parameterization for surface
511 3d radiative effects in large-eddy simulations. *Earth and Space Science Open Archive*, **17**.
512 <https://doi.org/10.1002/essoar.10511758>.
- 513 Veerman, M. A., X. Pedruzo-Bagazgoitia, F. Jakub, J. Vilà-Guerau de Arellano, C. C. van Heerwaarden,
514 2020: Three-Dimensional Radiative Effects By Shallow Cumulus Clouds on Dynamic Heterogenei-
515 ties Over a Vegetated Surface. *J. Advances Mod. Earth Sys.*, **12**, e2019MS00199,
516 <https://doi.org/10.1029/2019MS001990>
- 517 Wapler, K. and Mayer, B., 2008: A fast three-dimensional approximation for the calculation of surface
518 irradiance in Large-Eddy Simulation Models, *J. Appl. Meteorol. Clim.*, **47**, 3061–3071,
519 <https://doi.org/10.1175/2008JAMC1842.1>.
- 520 Welch, R. M and B. A. Wielicki, 1985: A radiative parameterization of stratocumulus cloud fields. *J.*
521 *Atmos. Sci.*, **42**, 2888-2897.
- 522